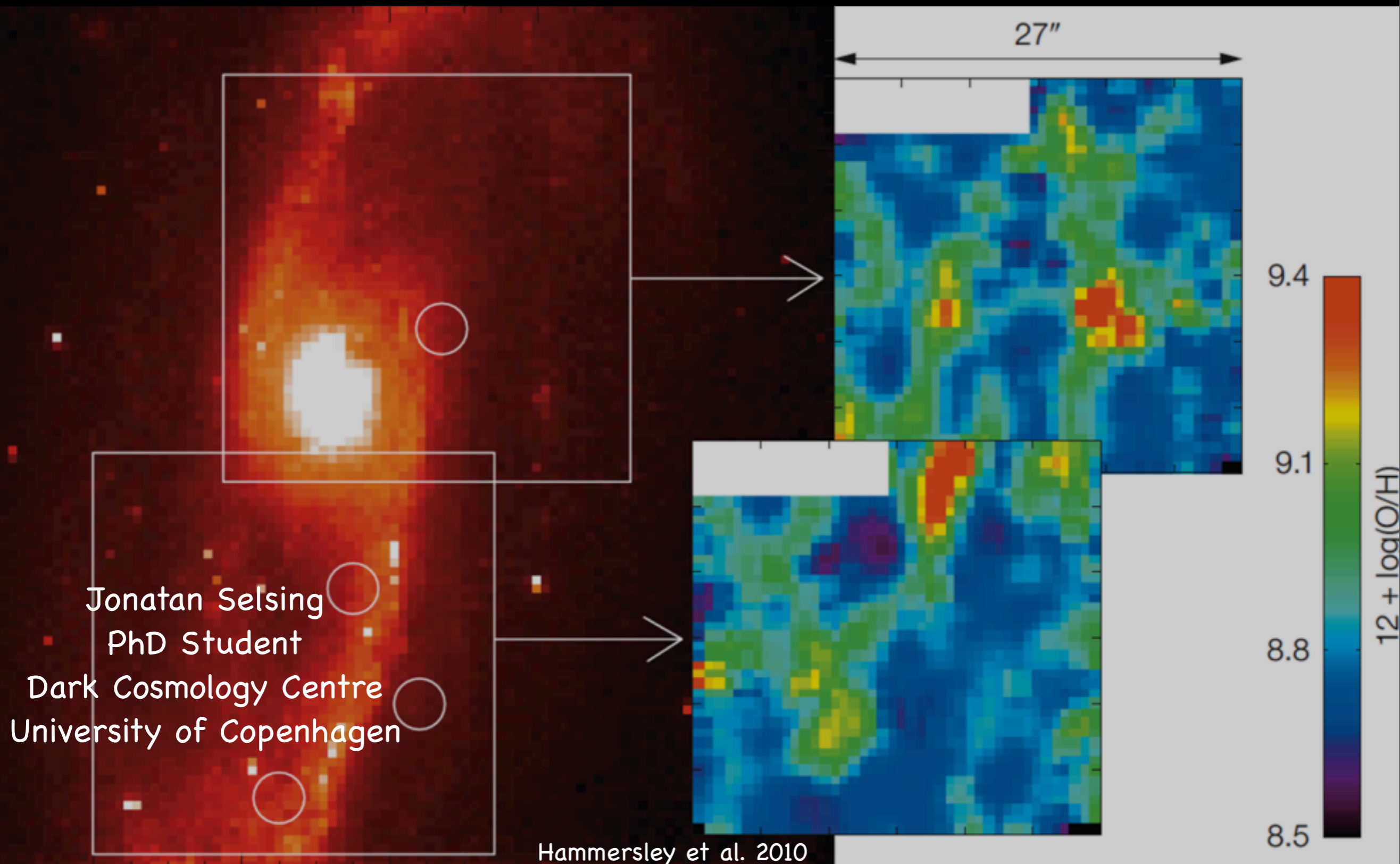


# Integral Field Spectroscopy of SN Ic/Ic - BL hosts

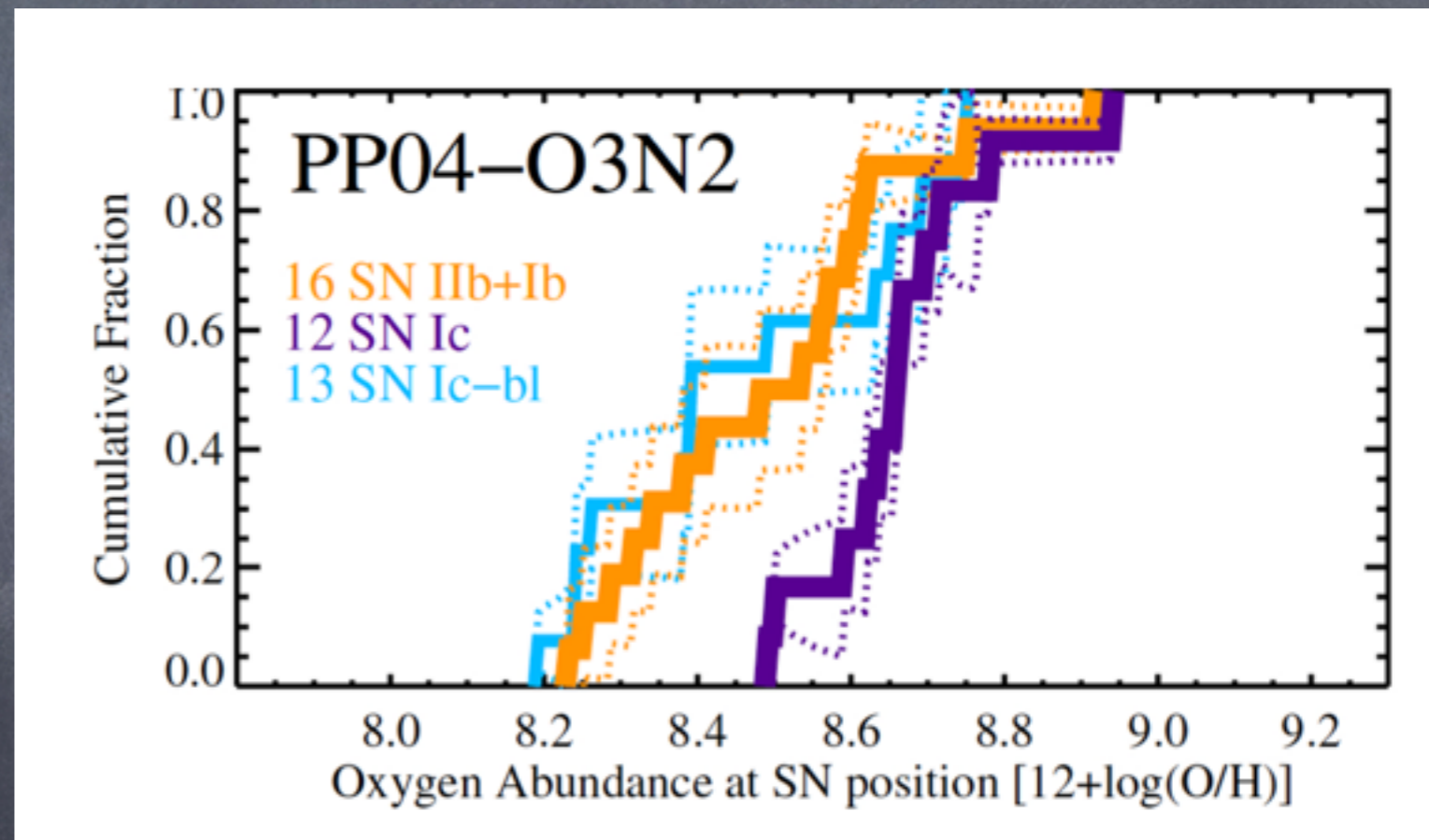


# Motivation

- What distinguishes a GRB progenitor from that of an ordinary SN Ic-bl without a GRB
- Investigate metallicity trends
- Constrain progenitor masses
- Study effect of local variations in host galaxy parameters

# Metallicity Trends

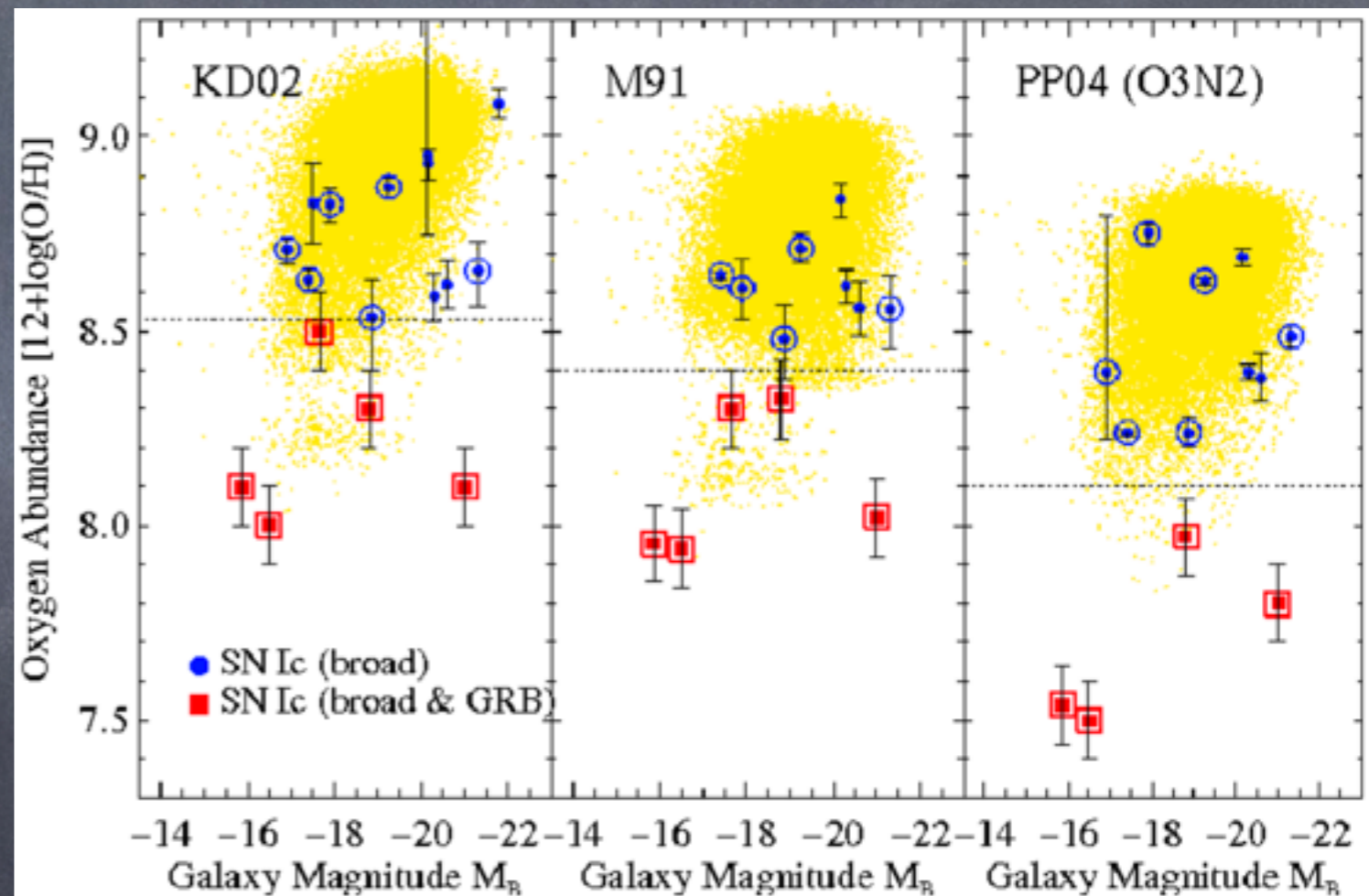
- SNe Type Ic lie in systematically more metal rich environments than other types of core collapse explosions



Modjaz, M. et al. 2011

# Metallicity Trends

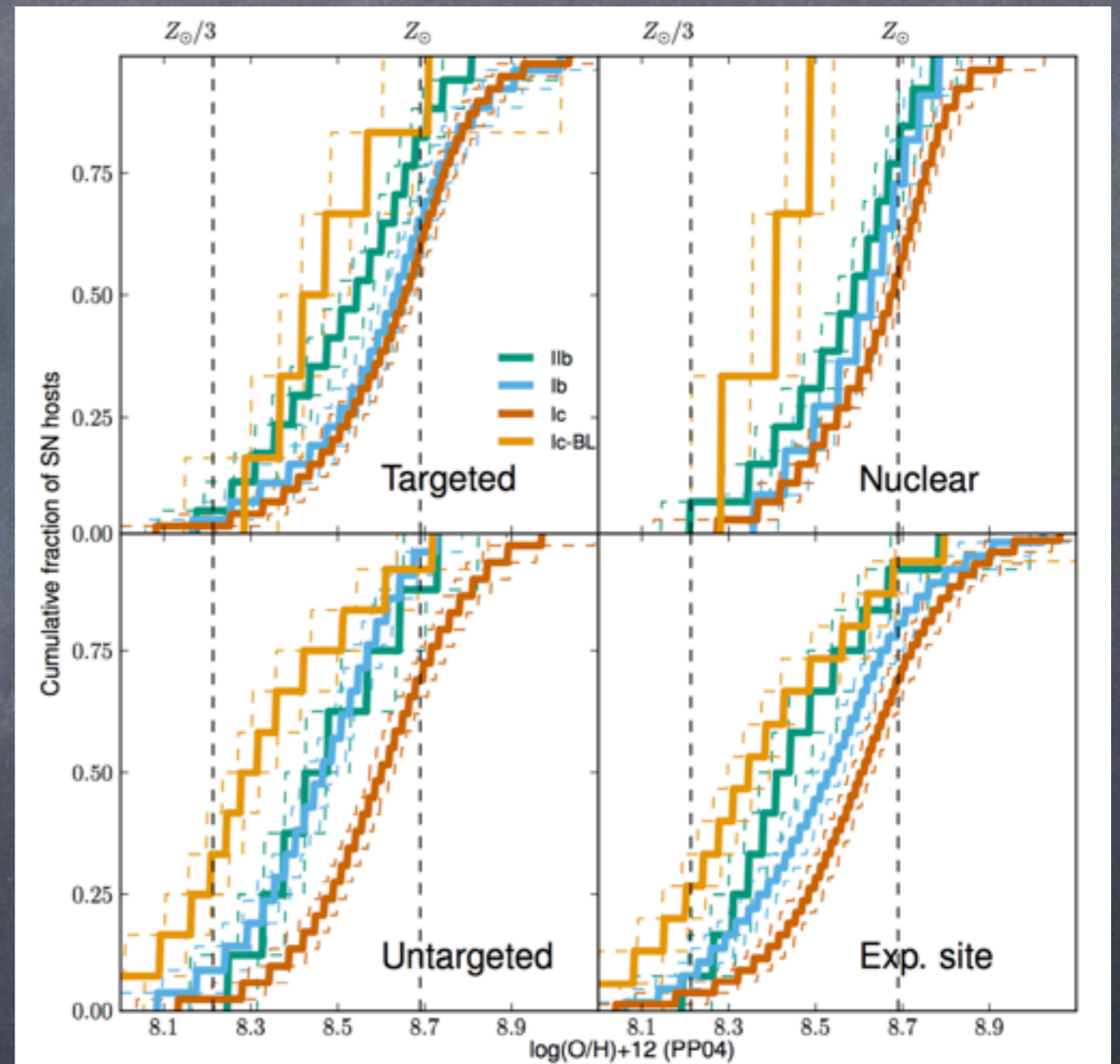
- SNe Type Ic lie in systematically more metal rich environments than other types of core collapse explosions
- SN Ic-bl without observed GRBs lie in systematically more metal rich environments than SNe with GRBs



Modjaz, M. et al. 2008

# Metallicity Trends

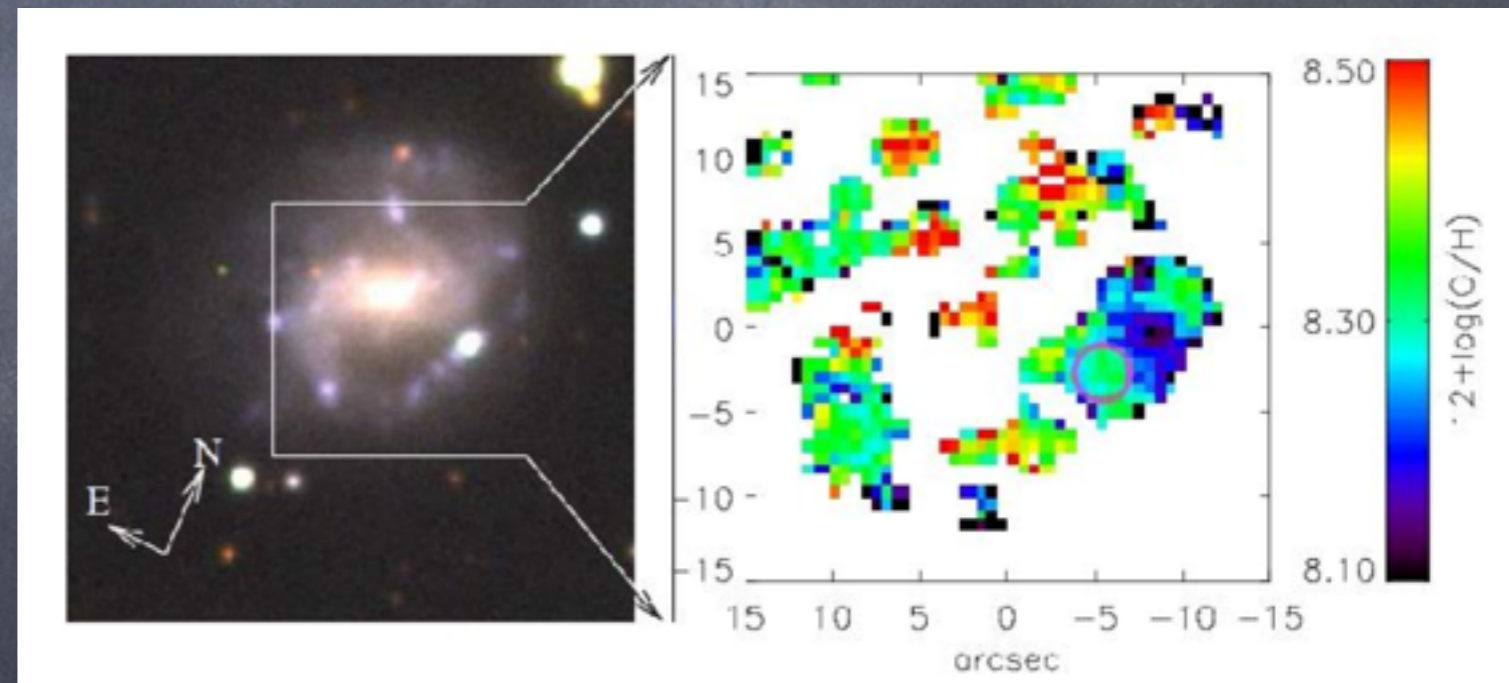
- SNe Type Ic lie in systematically more metal rich environments than other types of core collapse explosions
- SN Ic-bl without observed GRBs lie in systematically more metal rich environments than SNe with GRBs
- Biases can affect the metallicity trends investigated



Sanders et al. 2012

# Metallicity Trends

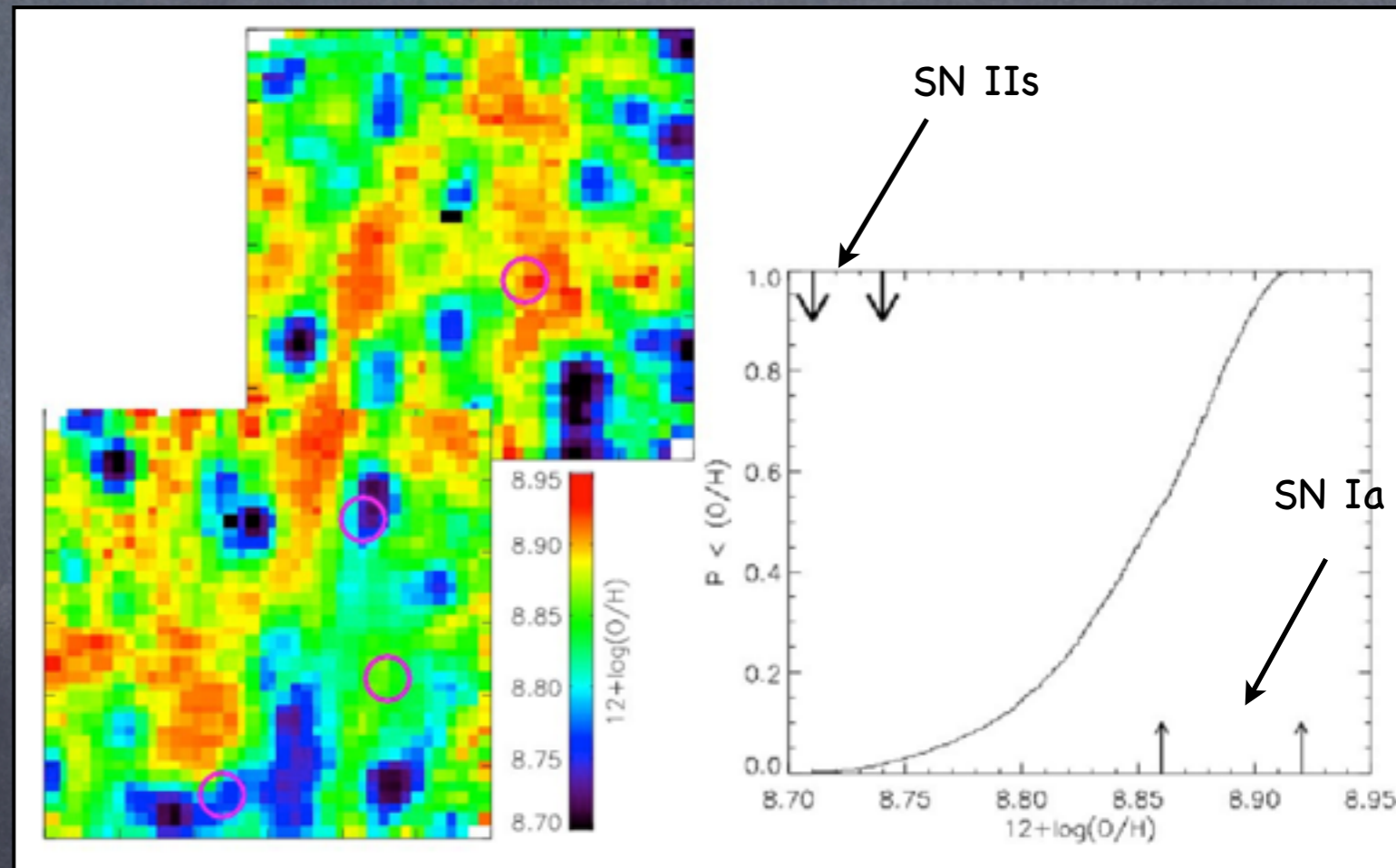
- SNe Type Ic lie in systematically more metal rich environments than other types of core collapse explosions
- SN Ic-bl without observed GRBs lie in systematically more metal rich environments than SNe with GRBs
- Biases can affect the metallicity trends investigated
- IFS is a possibility to probe local environments



Christensen, L. et al. 2008

# Metallicity Trends

- SNe Type Ic lie in systematically more metal rich environments than other types of core collapse explosions
- SN Ic-bl without observed GRBs lie in systematically more metal rich environments than SNe with GRBs
- Biases can affect the metallicity trends investigated
- IFS is a possibility to probe local environments
- Environmental differences for SNe



Hammersley et al. 2010

# Sample

- 16 Targets observed with VIMOS
- Host of both Ic (9) and Ic -BL (7) hosts
- Targeted (7) and non-targeted (9) hosts
- Spatially resolved - 250 pc resolution

## Hosts targeted

SN1996D T,Ic  
SN1997B T,Ic  
SN1999cn T,Ic  
SN2004fe T,Ic  
SN2006ck T,Ic  
SN2003jd T,Ic-BL  
SN1998ey T,Ic-BL  
SN2004bu T,Ic-BL  
SN2006ip non-T,Ic  
SN2007fj non-T,Ic  
SN2007hn non-T,Ic  
SN2009bh non-T,Ic  
SN2005ks non-T,Ic-BL  
SN2006qk non-T,Ic-BL  
SN2007I non-T,Ic-BL  
SN2007gx non-T,Ic-BL



# To be Measured Parameters

• Extinction  $\longrightarrow$  Balmer Decrement  
 $H\alpha/H\beta = 2.86$

# To be Measured Parameters

- Extinction

- Metallicity  $\longrightarrow$

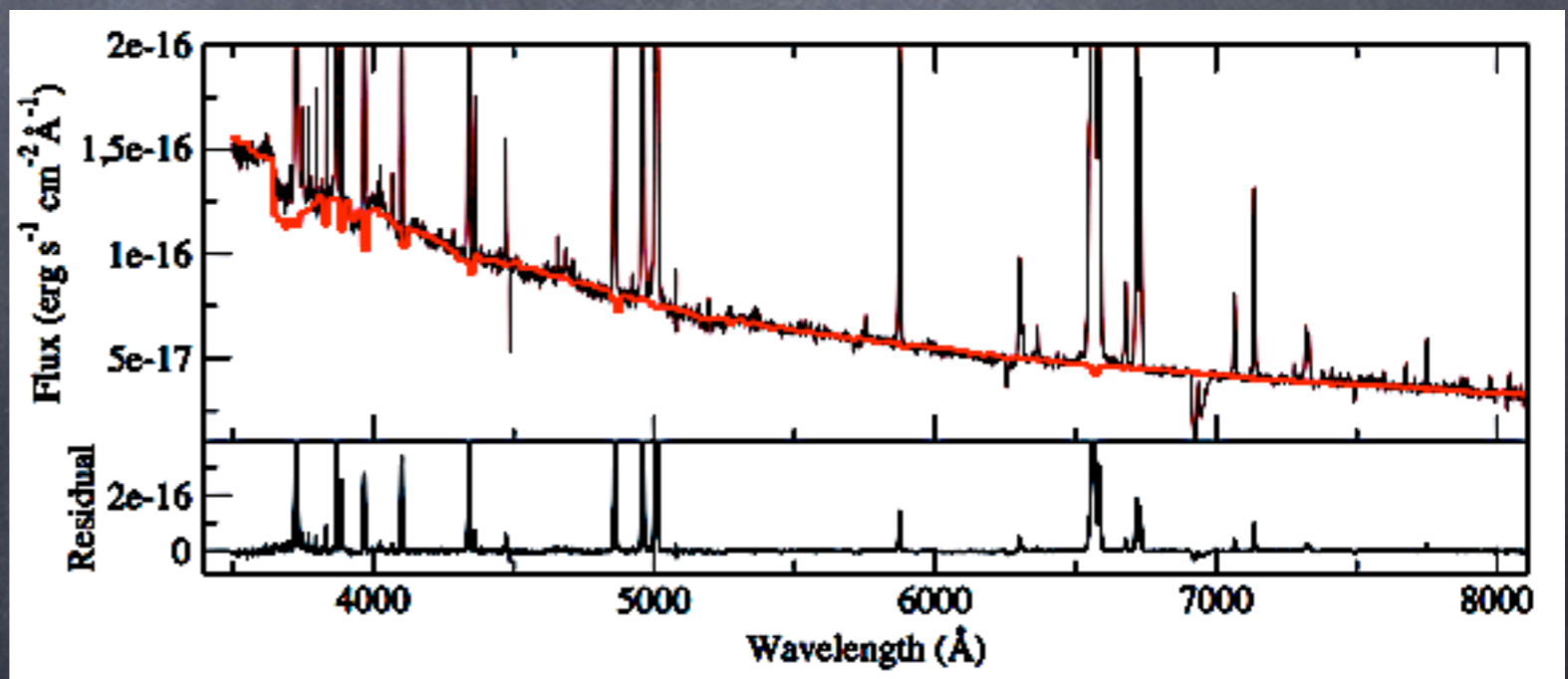
Strong Line Diagnostics

$$N2 : \log([Nii]/H\alpha)$$

$$O3N2 : \log\left(\frac{[Oiii]/H\beta}{[Nii]/H\alpha}\right)$$

# To be Measured Parameters

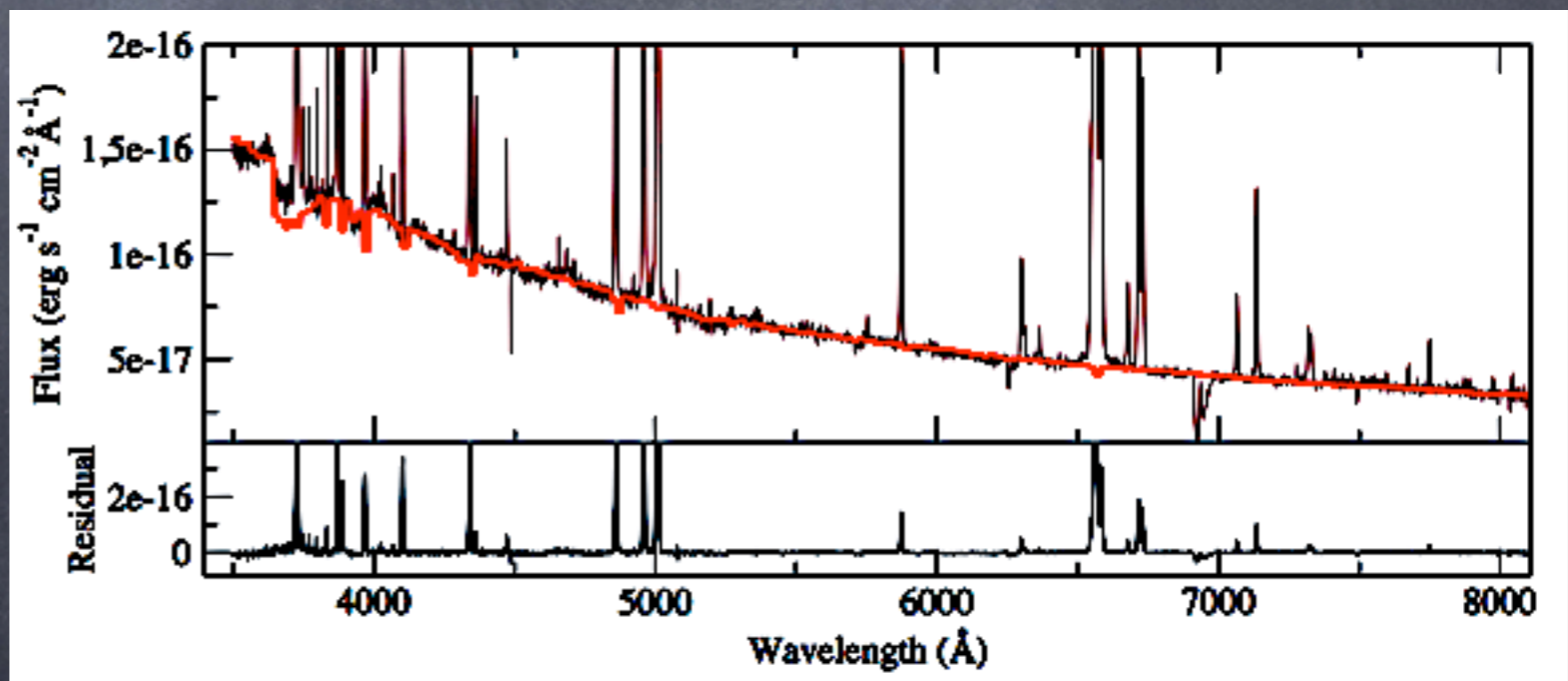
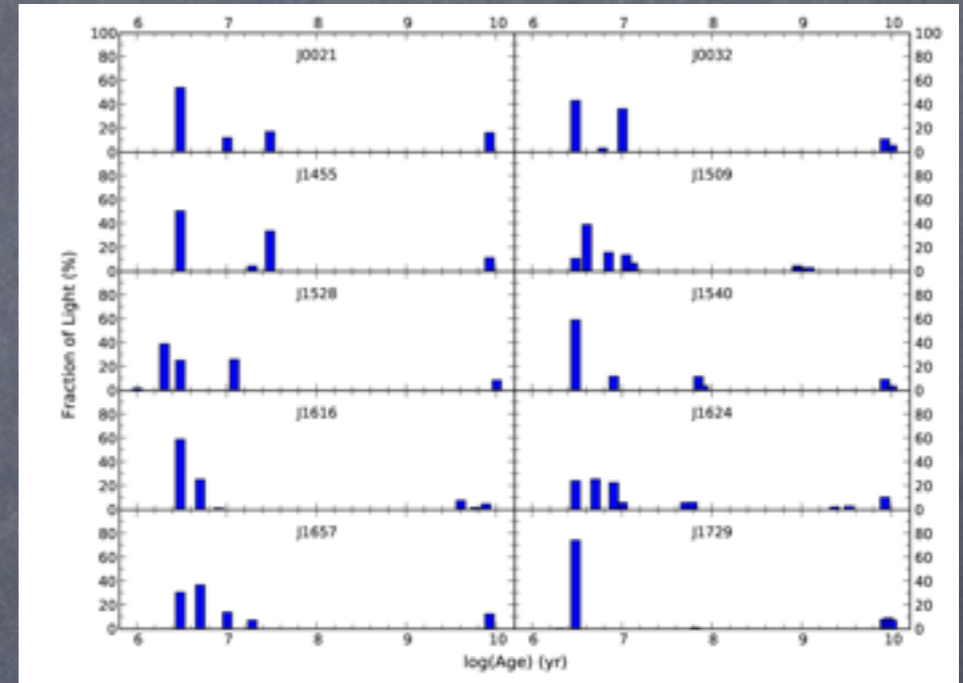
- Extinction
- Metallicity
- Stellar ages of the HII regions



E. Pérez-Montero et al. 2010

# To be Measured Parameters

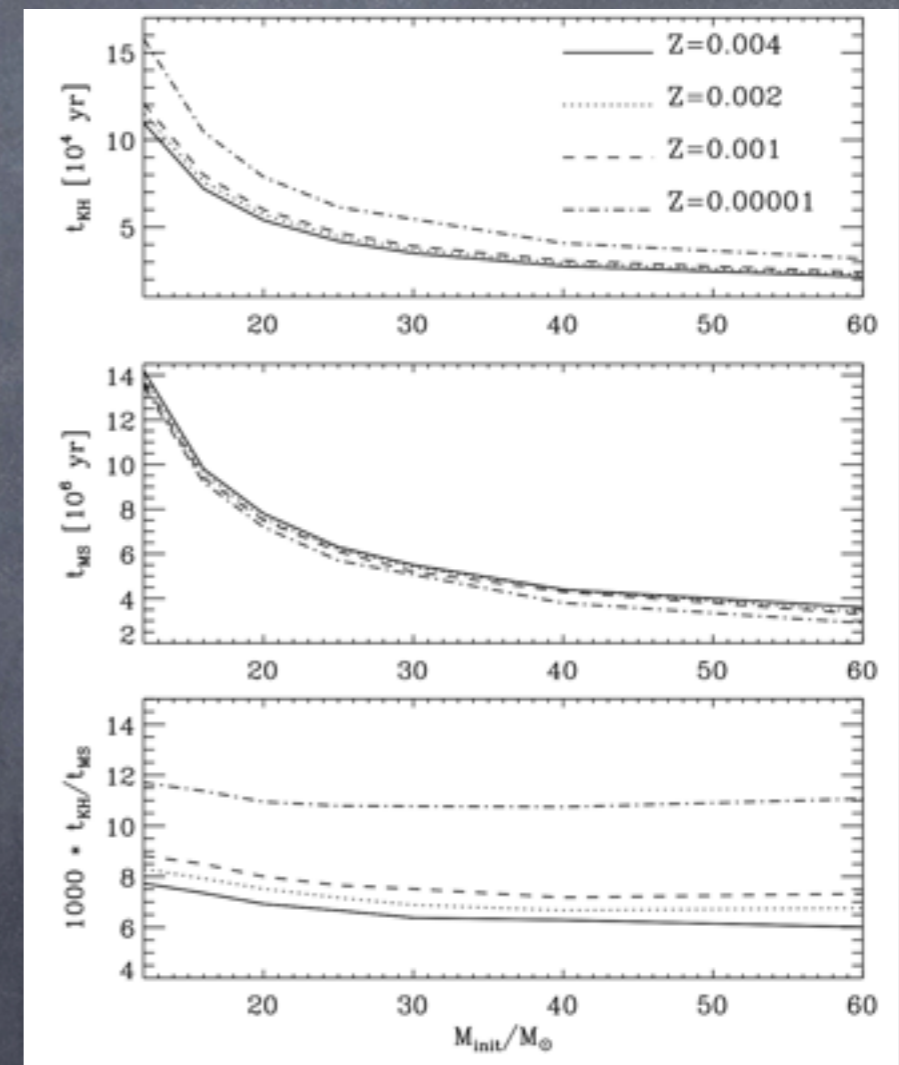
- Extinction
- Metallicity
- Stellar ages of the HII regions



E. Pérez-Montero et al. 2010

# To be Measured Parameters

- Extinction
- Metallicity
- Stellar ages of the HII regions
- Stellar masses of progenitors



S.-C. Yoon et al. 2006

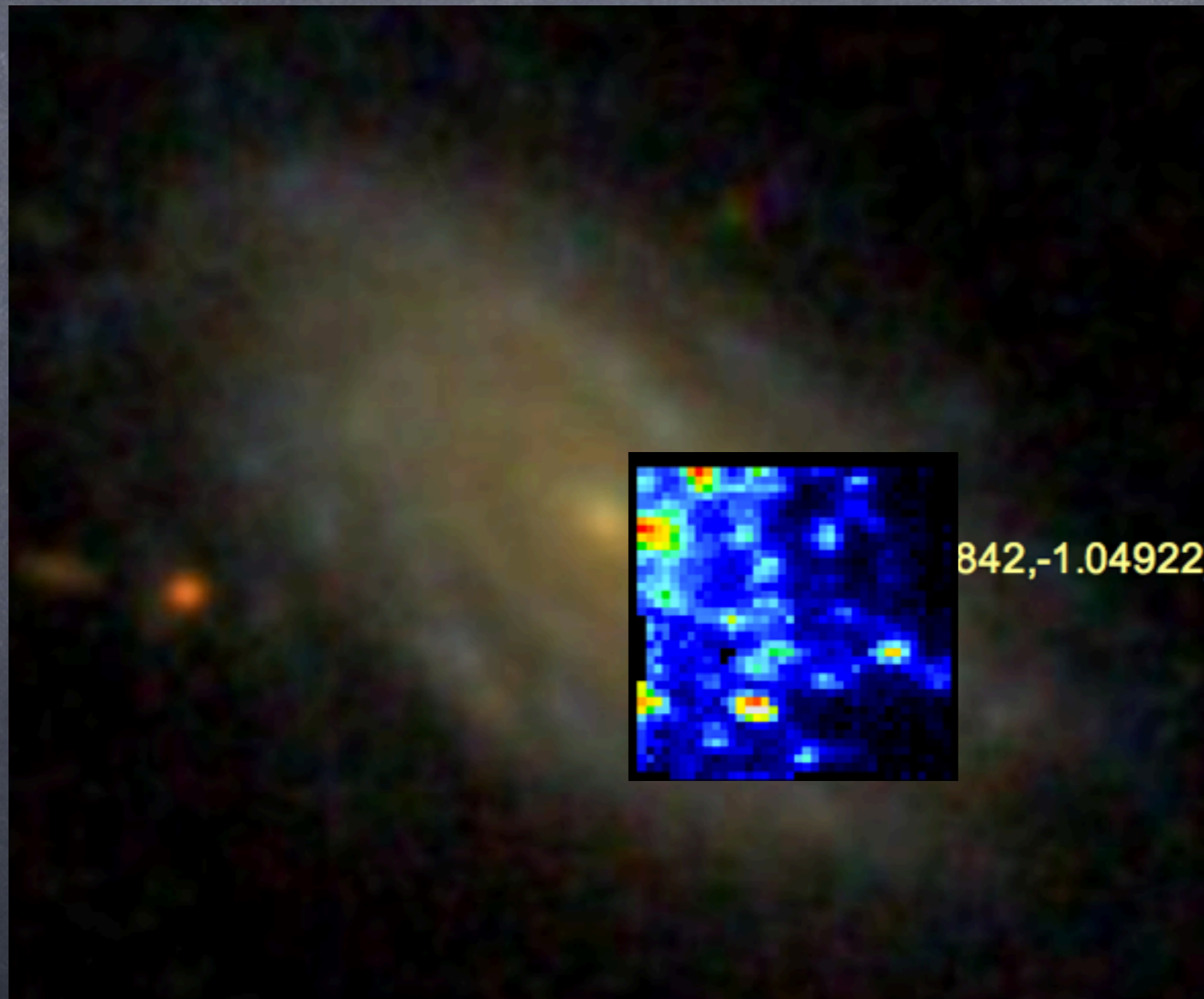
# A work in progress

SN2006ck(SN Ic) Host



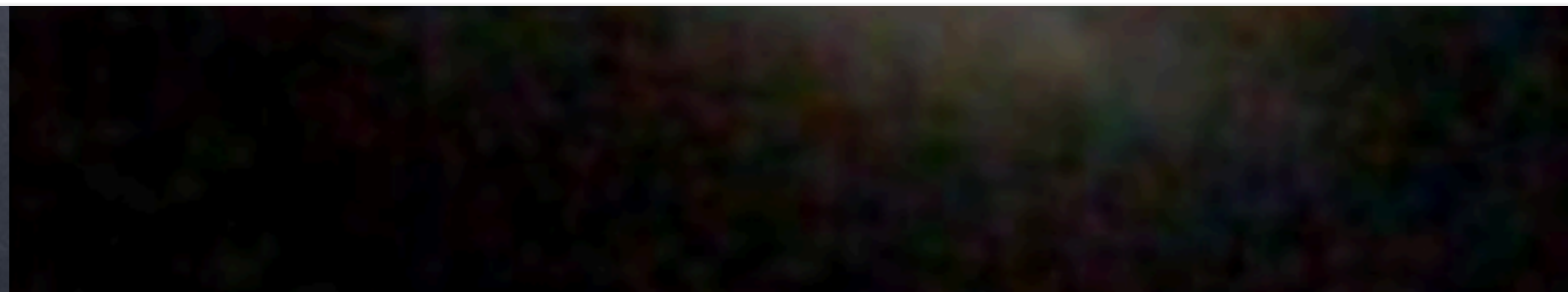
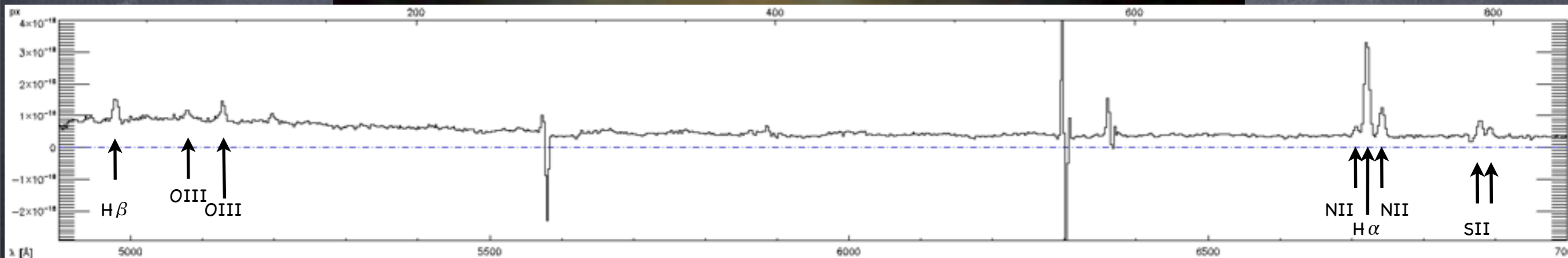
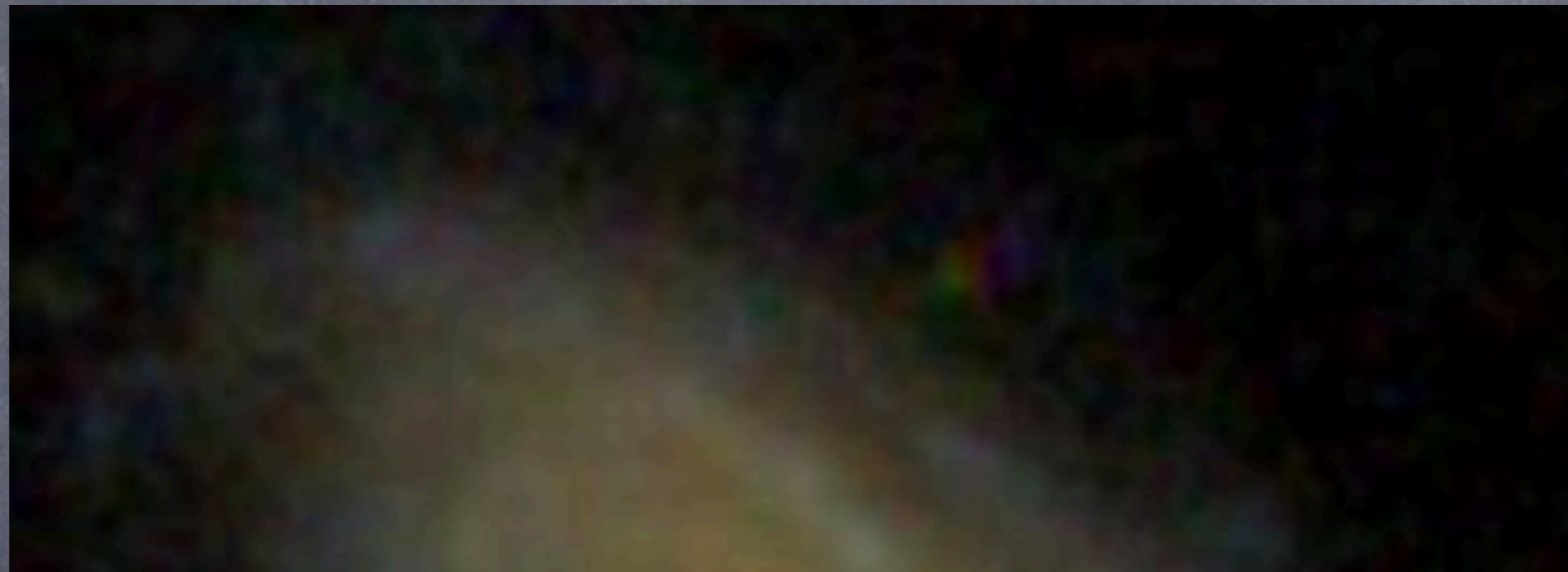
# A work in progress

SN2006ck(SN Ic) Host



# A work in progress

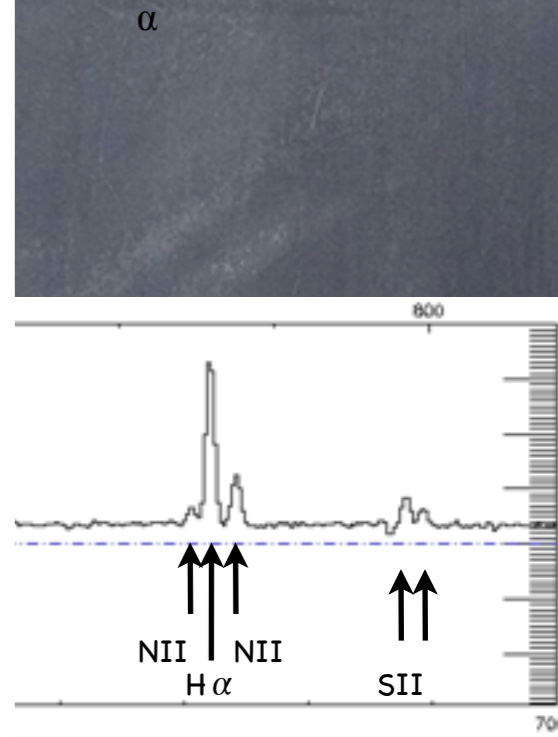
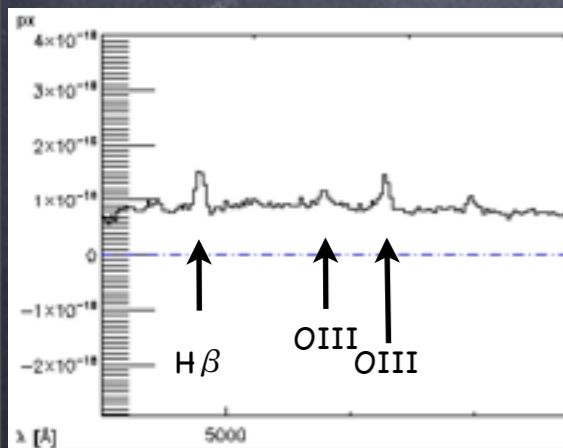
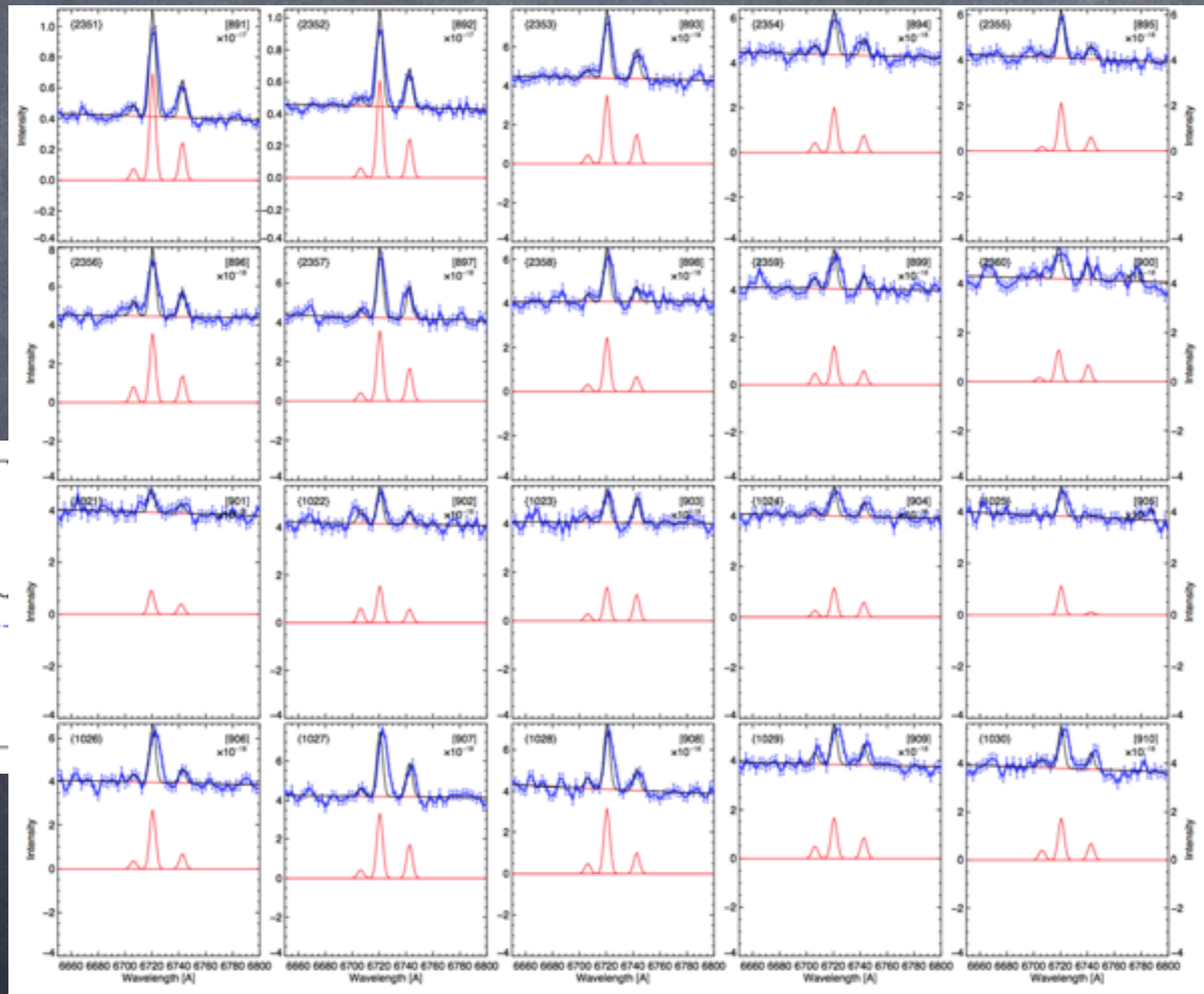
SN2006ck(SN Ic) Host





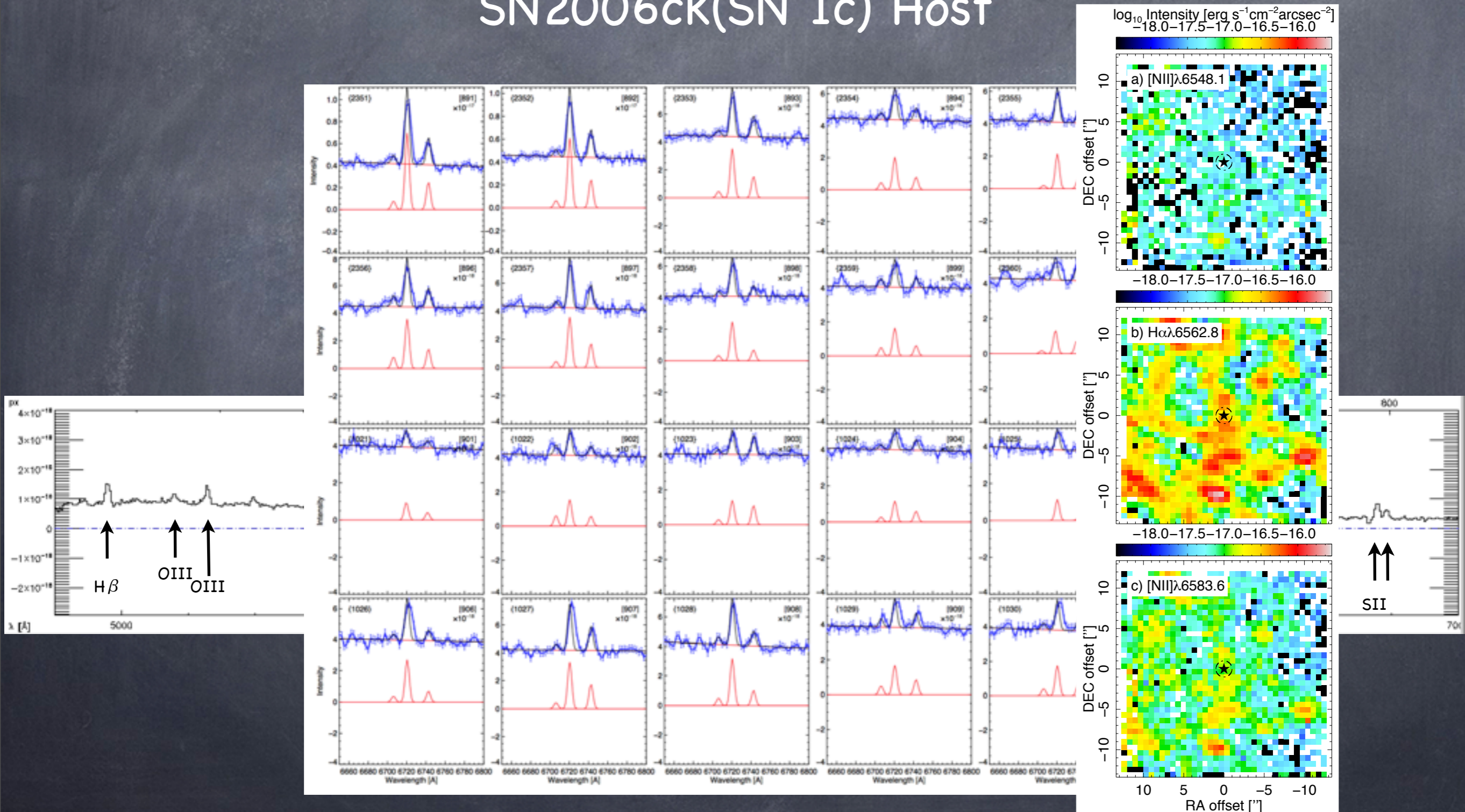
# A work in progress

SN2006ck(SN Ic) Host



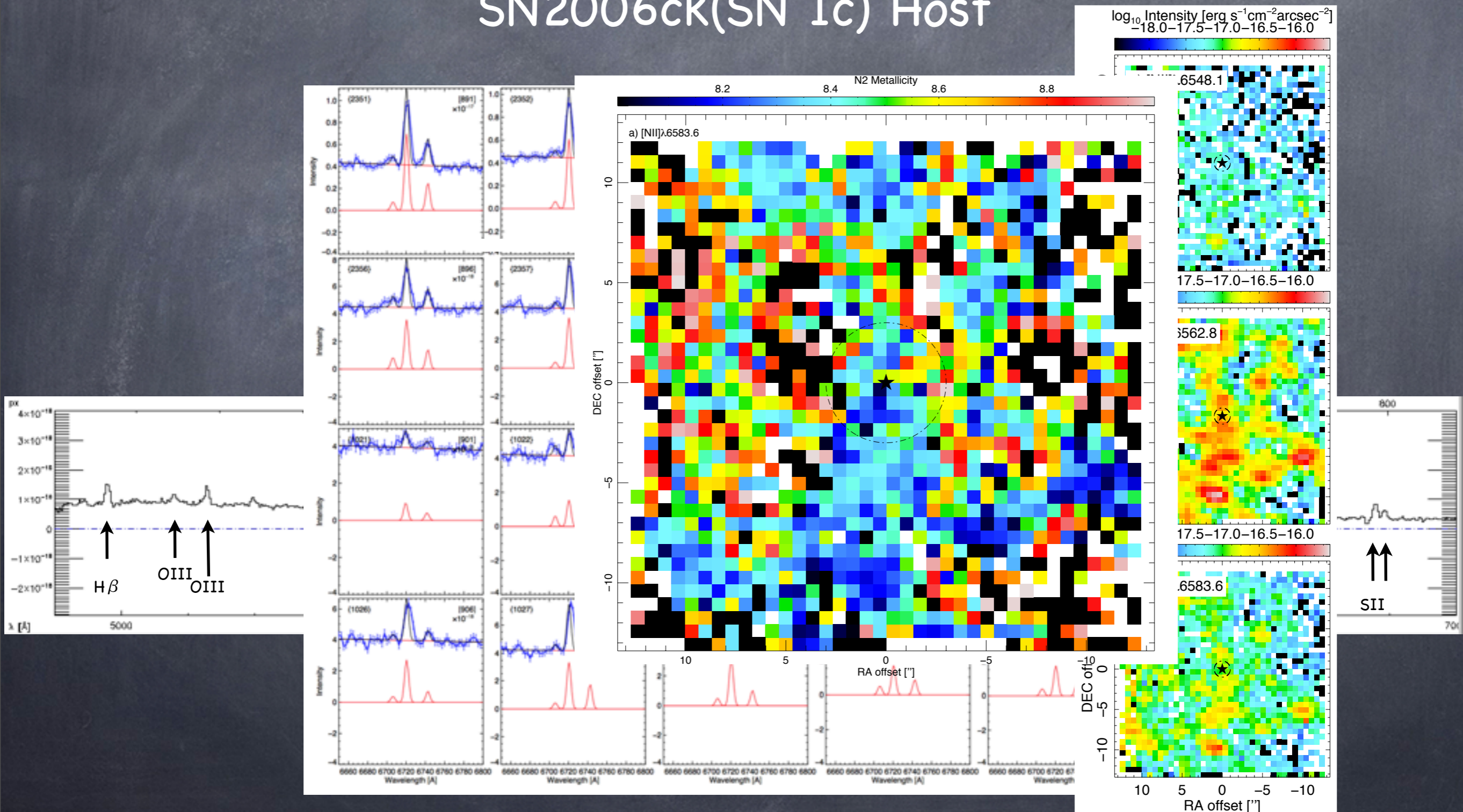
# A work in progress

## SN2006ck(SN Ic) Host



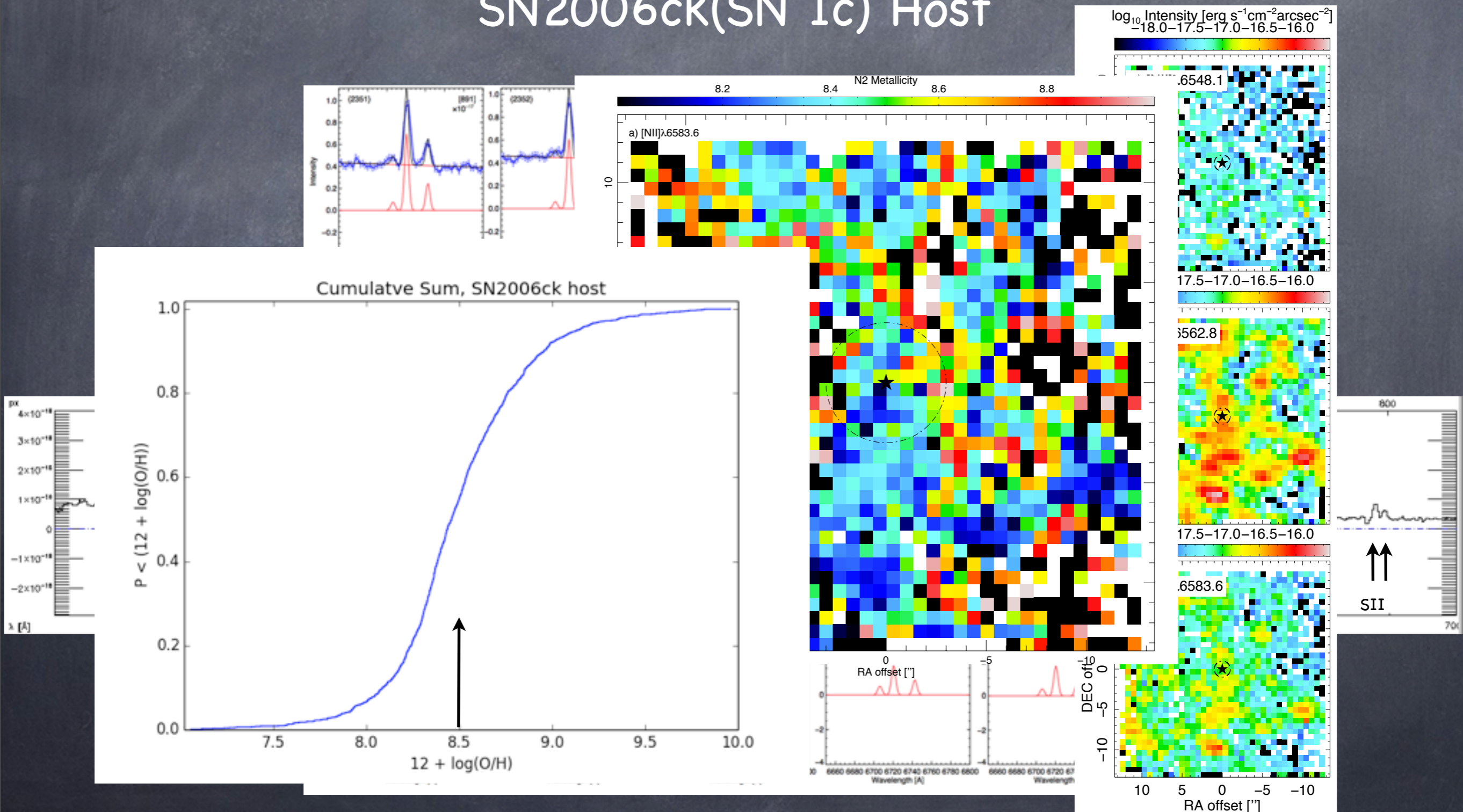
# A work in progress

## SN2006ck(SN Ic) Host



# A work in progress

## SN2006ck(SN Ic) Host



# Conclusions

👁 ?

👁 ?

👁 ?

👁 ?

👁 ?